## Search for Squarks and Gluinos in Events Containing Jets and a Large Imbalance in Transverse Energy

B. Abbott,<sup>40</sup> M. Abolins,<sup>37</sup> V. Abramov,<sup>15</sup> B.S. Acharya,<sup>8</sup> I. Adam,<sup>39</sup> D. L. Adams,<sup>49</sup> M. Adams,<sup>24</sup> S. Ahn,<sup>23</sup>
G. A. Alves,<sup>2</sup> N. Amos,<sup>36</sup> E. W. Anderson,<sup>30</sup> M. M. Baarmand,<sup>42</sup> V. V. Babintsev,<sup>15</sup> L. Babukhadia,<sup>16</sup> A. Baden,<sup>33</sup>
B. Baldin,<sup>23</sup> S. Banerjee,<sup>8</sup> J. Bantly,<sup>46</sup> E. Barberis,<sup>17</sup> P. Baringer,<sup>31</sup> J. F. Bartlett,<sup>23</sup> A. Belyaev,<sup>14</sup> S. B. Beri,<sup>6</sup> I. Bertram,<sup>26</sup> V. A. Bezzubov,<sup>15</sup> P. C. Bhat,<sup>23</sup> V. Bhatnagar,<sup>6</sup> M. Bhattacharjee,<sup>42</sup> N. Biswas,<sup>28</sup> G. Blazey,<sup>25</sup> I. Bertram,<sup>20</sup> V. A. Bezzubov,<sup>15</sup> P. C. Bhat,<sup>25</sup> V. Bhatnagar,<sup>6</sup> M. Bhattacharjee,<sup>42</sup> N. Biswas,<sup>20</sup> G. Blazey,<sup>25</sup>
S. Blessing,<sup>21</sup> P. Bloom,<sup>18</sup> A. Boehnlein,<sup>23</sup> N. I. Bojko,<sup>15</sup> F. Borcherding,<sup>23</sup> C. Boswell,<sup>20</sup> A. Brandt,<sup>23</sup> R. Breedon,<sup>18</sup> G. Briskin,<sup>46</sup> R. Brock,<sup>37</sup> A. Bross,<sup>23</sup> D. Buchholz,<sup>26</sup> V. S. Burtovoi,<sup>15</sup> J. M. Butler,<sup>34</sup> W. Carvalho,<sup>2</sup> D. Casey,<sup>37</sup> Z. Casilum,<sup>42</sup> H. Castilla-Valdez,<sup>11</sup> D. Chakraborty,<sup>42</sup> S. V. Chekulaev,<sup>15</sup> W. Chen,<sup>42</sup> S. Choi,<sup>10</sup> S. Chopra,<sup>21</sup> B. C. Choudhary,<sup>20</sup> J. H. Christenson,<sup>23</sup> M. Chung,<sup>24</sup> D. Claes,<sup>38</sup> A. R. Clark,<sup>17</sup> W. G. Cobau,<sup>33</sup> J. Cochran,<sup>20</sup> L. Coney,<sup>28</sup> W. E. Cooper,<sup>23</sup> D. Coppage,<sup>31</sup> C. Cretsinger,<sup>41</sup> D. Cullen-Vidal,<sup>46</sup> M. A. C. Cummings,<sup>25</sup> D. Cutts,<sup>46</sup> O. I. Dahl,<sup>17</sup> K. Davis,<sup>16</sup> K. De,<sup>47</sup> K. Del Signore,<sup>36</sup> M. Demarteau,<sup>23</sup> D. Denisov,<sup>23</sup> S. P. Denisov,<sup>15</sup> H. T. Diehl,<sup>23</sup> M. Diesburg,<sup>23</sup> G. Di Loreto,<sup>37</sup> P. Draper,<sup>47</sup> Y. Ducros,<sup>5</sup> L. V. Dudko,<sup>14</sup> S. R. Dugad,<sup>8</sup> A. Dyshkant,<sup>15</sup> D. Edmunds,<sup>37</sup> I. Ellison,<sup>20</sup> V. D. Elvira,<sup>42</sup> R. Engelmann,<sup>42</sup> S. Eno,<sup>33</sup> G. Enpley,<sup>49</sup> P. Ermoloy,<sup>14</sup> O. V. Eroshin,<sup>15</sup> V. N. Evdokimov,<sup>15</sup> J. Ellison,<sup>20</sup> V. D. Elvira,<sup>42</sup> R. Engelmann,<sup>42</sup> S. Eno,<sup>33</sup> G. Eppley,<sup>49</sup> P. Ermolov,<sup>14</sup> O. V. Eroshin,<sup>15</sup> V. N. Evdokimov,<sup>15</sup>
T. Fahland,<sup>19</sup> M. K. Fatyga,<sup>41</sup> S. Feher,<sup>23</sup> D. Fein,<sup>16</sup> T. Ferbel,<sup>41</sup> H. E. Fisk,<sup>23</sup> Y. Fisyak,<sup>43</sup> E. Flattum,<sup>23</sup> G. E. Forden,<sup>16</sup> M. Fortner,<sup>25</sup> K. C. Frame,<sup>37</sup> S. Fuess,<sup>23</sup> E. Gallas,<sup>47</sup> A. N. Galyaev,<sup>15</sup> P. Gartung,<sup>20</sup> V. Gavrilov,<sup>13</sup> T. L. Geld,<sup>37</sup> R. J. Genik II,<sup>37</sup> K. Genser,<sup>23</sup> C. E. Gerber,<sup>23</sup> Y. Gershtein,<sup>13</sup> B. Gibbard,<sup>43</sup> B. Gobbi,<sup>26</sup> B. Gómez,<sup>4</sup> G. Gómez,<sup>33</sup> R. J. Genik II,<sup>37</sup> K. Genser,<sup>25</sup> C. E. Gerber,<sup>25</sup> Y. Gershtein,<sup>15</sup> B. Gibbard,<sup>45</sup> B. Gobbi,<sup>26</sup> B. Gómez,<sup>4</sup> G. Gómez,<sup>55</sup> P. I. Goncharov,<sup>15</sup> J. L. González Solís,<sup>11</sup> H. Gordon,<sup>43</sup> L. T. Goss,<sup>48</sup> K. Gounder,<sup>20</sup> A. Goussiou,<sup>42</sup> N. Graf,<sup>43</sup> P. D. Grannis,<sup>42</sup> D. R. Green,<sup>23</sup> H. Greenlee,<sup>23</sup> S. Grinstein,<sup>1</sup> P. Grudberg,<sup>17</sup> S. Grünendahl,<sup>23</sup> G. Guglielmo,<sup>45</sup> J. A. Guida,<sup>16</sup> J. M. Guida,<sup>46</sup> A. Gupta,<sup>8</sup> S. N. Gurzhiev,<sup>15</sup> G. Gutierrez,<sup>23</sup> P. Gutierrez,<sup>45</sup> N. J. Hadley,<sup>33</sup>
H. Haggerty,<sup>23</sup> S. Hagopian,<sup>21</sup> V. Hagopian,<sup>21</sup> K. S. Hahn,<sup>41</sup> R. E. Hall,<sup>19</sup> P. Hanlet,<sup>35</sup> S. Hansen,<sup>23</sup> J. M. Hauptman,<sup>30</sup> C. Hebert,<sup>31</sup> D. Hedin,<sup>25</sup> A. P. Heinson,<sup>20</sup> U. Heintz,<sup>34</sup> R. Hernández-Montoya,<sup>11</sup> T. Heuring,<sup>21</sup> R. Hirosky,<sup>24</sup>
J. D. Hobbs,<sup>42</sup> B. Hoeneisen,<sup>4,\*</sup> J. S. Hoftun,<sup>46</sup> F. Hsieh,<sup>36</sup> Tong Hu,<sup>27</sup> A. S. Ito,<sup>23</sup> J. Jaques,<sup>28</sup> S. A. Jerger,<sup>37</sup> R. Jesik,<sup>27</sup> T. Joffe-Minor,<sup>26</sup> K. Johns,<sup>16</sup> M. Johnson,<sup>23</sup> A. Jonckheere,<sup>23</sup> M. Jones,<sup>22</sup> H. Jöstlein,<sup>23</sup> S. Y. Jun,<sup>26</sup> C. K. Jung,<sup>42</sup>
S. Kahn,<sup>43</sup> G. Kalbfleisch,<sup>45</sup> D. Karmanov,<sup>14</sup> D. Karmgard,<sup>21</sup> R. Kehoe,<sup>28</sup> S. K. Kim,<sup>10</sup> B. Klima,<sup>23</sup> C. Klopfenstein,<sup>18</sup> W. Ko<sup>18</sup> L. M. Kohli,<sup>6</sup> D. Koltick,<sup>29</sup> A. V. Kostritchiy,<sup>15</sup> L. Kotahar,<sup>43</sup> A. V. Kotural,<sup>39</sup> A. V. Kostrolov,<sup>15</sup> W. Ko,<sup>18</sup> J. M. Kohli,<sup>6</sup> D. Koltick,<sup>29</sup> A. V. Kostritskiy,<sup>15</sup> J. Kotcher,<sup>43</sup> A. V. Kotwal,<sup>39</sup> A. V. Kozelov,<sup>15</sup>
 E. A. Kozlovsky,<sup>15</sup> J. Krane,<sup>38</sup> M. R. Krishnaswamy,<sup>8</sup> S. Krzywdzinski,<sup>23</sup> S. Kuleshov,<sup>13</sup> Y. Kulik,<sup>42</sup> S. Kunori,<sup>33</sup> E. A. Kozlovsky,<sup>13</sup> J. Krane,<sup>36</sup> M. R. Krishnaswamy,<sup>6</sup> S. Krzywdzinski,<sup>25</sup> S. Kuleshov,<sup>15</sup> Y. Kulik,<sup>42</sup> S. Kunori,<sup>35</sup>
F. Landry,<sup>37</sup> G. Landsberg,<sup>46</sup> B. Lauer,<sup>30</sup> A. Leflat,<sup>14</sup> J. Li,<sup>47</sup> Q. Z. Li,<sup>23</sup> J. G. R. Lima,<sup>3</sup> D. Lincoln,<sup>23</sup> S. L. Linn,<sup>21</sup> J. Linnemann,<sup>37</sup> R. Lipton,<sup>23</sup> F. Lobkowicz,<sup>41</sup> A. Lucotte,<sup>42</sup> L. Lueking,<sup>23</sup> A. L. Lyon,<sup>33</sup> A. K. A. Maciel,<sup>2</sup>
R. J. Madaras,<sup>17</sup> R. Madden,<sup>21</sup> L. Magaña-Mendoza,<sup>11</sup> V. Manankov,<sup>14</sup> S. Mani,<sup>18</sup> H. S. Mao,<sup>23,†</sup> R. Markeloff,<sup>25</sup>
T. Marshall,<sup>27</sup> M. I. Martin,<sup>23</sup> K. M. Mauritz,<sup>30</sup> B. May,<sup>26</sup> A. A. Mayorov,<sup>15</sup> R. McCarthy,<sup>42</sup> J. McDonald,<sup>21</sup>
T. McKibben,<sup>24</sup> J. McKinley,<sup>37</sup> T. McMahon,<sup>44</sup> H. L. Melanson,<sup>23</sup> M. Merkin,<sup>14</sup> K. W. Merritt,<sup>23</sup> C. Miao,<sup>46</sup>
H. Miettinen,<sup>49</sup> A. Mincer,<sup>40</sup> C. S. Mishra,<sup>23</sup> N. Mokhov,<sup>23</sup> N. K. Mondal,<sup>8</sup> H. E. Montgomery,<sup>23</sup> P. Mooney,<sup>4</sup>
M. Mostafa,<sup>1</sup> H. da Motta,<sup>2</sup> C. Murphy,<sup>24</sup> F. Nang,<sup>16</sup> M. Narain,<sup>34</sup> V. S. Narasimham,<sup>8</sup> A. Narayanan,<sup>16</sup> H. A. Neal,<sup>36</sup>
J. P. Negret,<sup>4</sup> P. Nemethy,<sup>40</sup> D. Norman,<sup>48</sup> L. Oesch,<sup>36</sup> V. Oguri,<sup>3</sup> N. Oshima,<sup>23</sup> D. Owen,<sup>37</sup> P. Padley,<sup>49</sup> A. Para,<sup>23</sup> N. Parashar,<sup>35</sup> Y. M. Park,<sup>9</sup> R. Partridge,<sup>46</sup> N. Parua,<sup>8</sup> M. Paterno,<sup>41</sup> B. Pawlik,<sup>12</sup> J. Perkins,<sup>47</sup> M. Peters,<sup>22</sup> R. Piegaia,<sup>1</sup> H. Piekarz,<sup>21</sup> Y. Pischalnikov,<sup>29</sup> B. G. Pope,<sup>37</sup> H. B. Prosper,<sup>21</sup> S. Protopopescu,<sup>43</sup> J. Qian,<sup>36</sup> P. Z. Quintas,<sup>23</sup> R. Raja,<sup>23</sup> S. Rajagopalan,<sup>43</sup> O. Ramirez,<sup>24</sup> S. Reucroft,<sup>35</sup> M. Rijssenbeek,<sup>42</sup> T. Rockwell,<sup>37</sup> M. Roco,<sup>23</sup> P. Rubinov,<sup>26</sup> R. Ruchti,<sup>28</sup> J. Rutherfoord,<sup>16</sup> A. Sánchez-Hernández,<sup>11</sup> A. Santoro,<sup>2</sup> L. Sawyer,<sup>32</sup> R. D. Schamberger,<sup>42</sup> H. Schellman,<sup>26</sup> J. Sculli,<sup>40</sup> E. Shabalina,<sup>14</sup> C. Shaffer,<sup>21</sup> H. C. Shankar,<sup>8</sup> R. K. Shivpuri,<sup>7</sup> D. Shpakov,<sup>42</sup> M. Shupe,<sup>16</sup> H. Singh,<sup>20</sup> J. B. Singh,<sup>6</sup> V. Sirotenko,<sup>25</sup> E. Smith,<sup>45</sup> R. P. Smith,<sup>23</sup> R. Snihur,<sup>26</sup> G. R. Snow,<sup>38</sup> J. Snow,<sup>44</sup> S. Snyder,<sup>43</sup> J. Solomon,<sup>24</sup>
 M. Sosebee,<sup>47</sup> N. Sotnikova,<sup>14</sup> M. Souza,<sup>2</sup> G. Steinbrück,<sup>45</sup> R. W. Stephens,<sup>47</sup> M. L. Stevenson,<sup>17</sup> F. Stichelbaut,<sup>43</sup> M. Sosebee,<sup>47</sup> N. Sotnikova,<sup>14</sup> M. Souza,<sup>2</sup> G. Steinbrück,<sup>43</sup> R. W. Stephens,<sup>47</sup> M. L. Stevenson,<sup>17</sup> F. Stichelbaut,<sup>43</sup> D. Stoker,<sup>19</sup> V. Stolin,<sup>13</sup> D. A. Stoyanova,<sup>15</sup> M. Strauss,<sup>45</sup> K. Streets,<sup>40</sup> M. Strovink,<sup>17</sup> A. Sznajder,<sup>2</sup> P. Tamburello,<sup>33</sup> J. Tarazi,<sup>19</sup> M. Tartaglia,<sup>23</sup> T. L. T. Thomas,<sup>26</sup> J. Thompson,<sup>33</sup> T. G. Trippe,<sup>17</sup> P. M. Tuts,<sup>39</sup> V. Vaniev,<sup>15</sup> N. Varelas,<sup>24</sup> E. W. Varnes,<sup>17</sup> A. A. Volkov,<sup>15</sup> A. P. Vorobiev,<sup>15</sup> H. D. Wahl,<sup>21</sup> G. Wang,<sup>21</sup> J. Warchol,<sup>28</sup> G. Watts,<sup>46</sup> M. Wayne,<sup>28</sup> H. Weerts,<sup>37</sup> A. White,<sup>47</sup> J. T. White,<sup>48</sup> J. A. Wightman,<sup>30</sup> S. Willis,<sup>25</sup> S. J. Wimpenny,<sup>20</sup> J. V. D. Wirjawan,<sup>48</sup> J. Womersley,<sup>23</sup> E. Won,<sup>41</sup> D. R. Wood,<sup>35</sup> Z. Wu,<sup>23,†</sup> R. Yamada,<sup>23</sup> P. Yamin,<sup>43</sup> T. Yasuda,<sup>35</sup> P. Yepes,<sup>49</sup> K. Yip,<sup>23</sup> C. Yoshikawa,<sup>22</sup> S. Youssef,<sup>21</sup> J. Yu,<sup>23</sup> Y. Yu,<sup>10</sup> B. Zhang,<sup>23,†</sup> Z. Zhou,<sup>30</sup> Z. H. Zhu,<sup>41</sup> M. Zielinski,<sup>41</sup> D. Zieminska,<sup>27</sup> A. Zieminski,<sup>27</sup> E. G. Zverev,<sup>14</sup> and A. Zylberstejn<sup>5</sup>

(D0 Collaboration)

<sup>1</sup>Universidad de Buenos Aires, Buenos Aires, Argentina <sup>2</sup>LAFEX, Centro Brasileiro de Pesquisas Físicas, Rio de Janiero, Brazil

<sup>3</sup>Universidade do Estado do Rio de Janeiro, Rio de Janeiro, Brazil

0031-9007/99/83(24)/4937(6)\$15.00 © 1999 The American Physical Society

<sup>4</sup>Universidad de los Andes, Bogotá, Colombia <sup>5</sup>DAPNIA/Service de Physique des Particules, CEA, Saclay, France <sup>6</sup>Panjab University, Chandigarh, India <sup>7</sup>Delhi University, Delhi, India <sup>8</sup>Tata Institute of Fundamental Research, Mumbai, India <sup>9</sup>Kyungsung University, Pusan, Korea <sup>10</sup>Seoul National University, Seoul, Korea <sup>11</sup>CINVESTAV, Mexico City, Mexico <sup>12</sup>Institute of Nuclear Physics, Kraków, Poland <sup>13</sup>Institute for Theoretical and Experimental Physics, Moscow, Russia <sup>14</sup>Moscow State University, Moscow, Russia <sup>15</sup>Institute for High Energy Physics, Protvino, Russia <sup>16</sup>University of Arizona, Tucson, Arizona 85721 <sup>17</sup>Lawrence Berkeley National Laboratory and University of California, Berkeley, California 94720 <sup>18</sup>University of California, Davis, California 95616 <sup>19</sup>University of California, Irvine, California 92697 <sup>20</sup>University of California, Riverside, California 92521 <sup>21</sup>Florida State University, Tallahassee, Florida 32306 <sup>22</sup>University of Hawaii, Honolulu, Hawaii 96822 <sup>23</sup>Fermi National Accelerator Laboratory, Batavia, Illinois 60510 <sup>24</sup>University of Illinois at Chicago, Chicago, Illinois 60607 <sup>25</sup>Northern Illinois University, DeKalb, Illinois 60115 <sup>26</sup>Northwestern University, Evanston, Illinois 60208 <sup>27</sup>Indiana University, Bloomington, Indiana 47405 <sup>28</sup>University of Notre Dame, Notre Dame, Indiana 46556 <sup>29</sup>Purdue University, West Lafayette, Indiana 47907 <sup>30</sup>Iowa State University, Ames, Iowa 50011 <sup>31</sup>University of Kansas, Lawrence, Kansas 66045 <sup>32</sup>Louisiana Tech University, Ruston, Louisiana 71272 <sup>33</sup>University of Maryland, College Park, Maryland 20742 <sup>34</sup>Boston University, Boston, Massachusetts 02215 <sup>35</sup>Northeastern University, Boston, Massachusetts 02115 <sup>36</sup>University of Michigan, Ann Arbor, Michigan 48109 <sup>37</sup>Michigan State University, East Lansing, Michigan 48824 <sup>38</sup>University of Nebraska, Lincoln, Nebraska 68588 <sup>39</sup>Columbia University, New York, New York 10027 <sup>40</sup>New York University, New York, New York 10003 <sup>41</sup>University of Rochester, Rochester, New York 14627 <sup>42</sup>State University of New York, Stony Brook, New York 11794 <sup>43</sup>Brookhaven National Laboratory, Upton, New York 11973 <sup>44</sup>Langston University, Langston, Oklahoma 73050 <sup>45</sup>University of Oklahoma, Norman, Oklahoma 73019 <sup>46</sup>Brown University, Providence, Rhode Island 02912 <sup>47</sup>University of Texas, Arlington, Texas 76019 <sup>48</sup>Texas A&M University, College Station, Texas 77843 <sup>49</sup>Rice University, Houston, Texas 77005

(Received 10 February 1999; revised manuscript received 14 September 1999)

Using data corresponding to an integrated luminosity of 79 pb<sup>-1</sup>, D0 has searched for events containing multiple jets and large missing transverse energy in  $p\overline{p}$  collisions at  $\sqrt{s} = 1.8$  TeV at the Fermilab Tevatron collider. Observing no significant excess beyond what is expected from the standard model, we set limits on the masses of squarks and gluinos and on the model parameters  $m_0$  and  $m_{1/2}$ , in the framework of the minimal low-energy supergravity models of supersymmetry. For  $\tan \beta = 2$  and  $A_0 = 0$ , with  $\mu < 0$ , we exclude all models with  $m_{\tilde{q}} < 250 \text{ GeV}/c^2$ . For models with equal squark and gluino masses, we exclude  $m < 260 \text{ GeV}/c^2$ .

PACS numbers: 12.60.Jv, 13.85.Rm, 14.80.Ly

Supersymmetry (SUSY) [1] is a symmetry that relates fermions and bosons, and can solve the hierarchy problem of the Higgs sector of the standard model (SM) [2]. Minimal SUSY extensions of the SM (MSSM) require partners (sparticles) for all standard model particles: a scalar partner for each quark and lepton (called squarks and sleptons), and a spin-half partner for each of the gauge bosons and Higgs scalars, which form the gluinos and

the mixed states called charginos and neutralinos. Such models also require four Higgs particles. Each particle in a SUSY model has an internal quantum number called Rparity. If R is conserved, as is assumed in this analysis, then sparticle states must be produced in pairs, and each sparticle that decays must contain an odd number of sparticles in its decay products. Consequently, the lightest SUSY particle (LSP) must be stable.

Because the most general supersymmetric extension of the SM has over 100 undetermined parameters, models have been developed that contain additional symmetries and constraints. Here we consider gravity-mediated SUSY breaking models, called minimal low-energy supergravity (mSUGRA) [3], where the scalar (squark and slepton) masses are unified to a single value  $m_0$  at the grand unified theory energy scale, and the gaugino masses are unified to a single value  $m_{1/2}$ . Three other parameters describe the Higgs and gaugino sectors of the model:  $\tan\beta$ , the ratio of the vacuum expectation values of the two Higgs doublets;  $A_0$ , a universal trilinear coupling constant; and the sign of  $\mu$ , a mixing parameter in the Higgsino mass matrix. For models in which the lightest neutralino  $(\tilde{\chi}_1^0)$  is the LSP, the LSP interacts only weakly and therefore cannot be observed directly, providing an excellent experimental SUSY signature: large missing transverse energy  $(\not\!\!E_T)$ . In such models, squarks  $(\tilde{q})$  and gluinos  $(\tilde{g})$ can decay through a cascade of charginos and neutralinos to final states consisting of quarks, leptons, and the LSP. Here we describe a search for squarks and gluinos in the jets and  $\not\!\!E_T$  channel.

The data, corresponding to an integrated luminosity of  $79.2 \pm 4.2 \text{ pb}^{-1}$ , were collected with the D0 detector [4] at the Fermilab Tevatron  $p\overline{p}$  collider operating at a centerof-mass energy of 1.8 TeV during 1993–1995. D0 has three major components: a central tracking system, central and forward uranium/liquid-argon calorimeters with towers in pseudorapidity and azimuth of  $\Delta \eta \times \Delta \phi = 0.1 \times 0.1$ , and a toroidal muon spectrometer. Jets are reconstructed using a cone algorithm [5] with a cone radius of 0.5 in  $\eta$ - $\phi$  space. The electromagnetic energy scale is set using the  $Z \rightarrow ee$  signal. The jet energy scale is determined from energy balance in events containing a hadronic jet and a photon candidate. The  $\not{E}_T$  is calculated from the vector sum of energy deposited in all calorimeter cells.

The initial data set was collected using an on-line hardware trigger that required  $\not\!\!E_T > 40$  GeV and at least one calorimeter trigger tower (of size  $\Delta \eta \times \Delta \phi = 0.2 \times$ 0.2) with transverse energy  $E_T > 5$  GeV. Furthermore, a software off-line filter required that events have  $\not\!\!E_T >$ 40 GeV and at least two jets with  $E_T > 8$  GeV.

To remove events with false large  $\not E_T$  due to detector noise and losses from the accelerator, we required events to have a summed scalar  $E_T$  ( $S_T$ ),  $0.0 < S_T < 1.8$  TeV. The position of the primary interaction vertex is also required to be within 60 cm of the detector center. This initial data sample contains 71 023 events. We required that all jets in the event with  $E_T > 15$  GeV meet quality criteria based on cluster shape [6], and that the three jets with the highest  $E_T$  be within  $|\eta| < 1.1$ , or within  $1.4 < |\eta| < 3.5$ . The shape criteria reject events with large  $\not{E}_T$  caused by poorly measured jets and detector noise and events where a jet deposited more than 90% of its energy in the electromagnetic portion of the calorimeter. Events with real electrons, such as production of  $W \rightarrow e\nu$  with jets, are thus effectively eliminated.

To select events consistent with the signal, we required at least three jets with  $E_T > 25$  GeV. In order to use a jet trigger for background studies, we accepted only events where the leading jet had  $E_T > 115$  GeV. 2723 events remain in the data at this point. We required at least  $\not{E}_T > 75$  GeV ( $\not{E}_T$  threshold varies with signal sample) in order to be in the region where our analysis trigger was fully efficient. These requirements leave 544 events.

To suppress quantum chromodynamics (QCD) multijet background, we required the azimuthal difference between the  $\not{E}_T$  and a jet of  $E_T > 25$  GeV be  $\delta \phi > 0.1$ , or  $<(\pi - 0.1)$  radians. We also required  $(\delta \phi_1 - \pi)^2 + \delta \phi_2^2 \ge$  $(0.5)^2$ , where 1 (2) denotes the leading (second-leading) jet in  $E_T$ , to reject events where a fluctuation of the second leading jet masks a fluctuation of the leading jet. 87 events remain. To reduce the background from *W* and *Z* boson production in association with jets, we required at least  $H_T > 100$  GeV ( $H_T$  threshold varies with signal sample), where  $H_T$  is defined as the scalar sum of the transverse energies of all but the leading jet. 50 events remain. To remove the remaining  $W \rightarrow \mu \nu +$  jets events, we rejected events containing isolated muons with  $p_T > 15$  GeV/*c*. A total of 49 events remain.

The average Tevatron luminosity was  $\approx 9 \times 10^{30} \text{ cm}^{-2} \text{ s}^{-1}$ , and peaked at about  $2 \times 10^{31} \text{ cm}^{-2} \text{ s}^{-1}$ . At the average, there is a 75% probability of having an additional  $p\overline{p}$  interaction accompanying the hard scattering. These additional events contribute many charged tracks, which can occasionally cause the soft collision to be chosen as the primary interaction vertex and result in a gross mismeasurement of the  $\not{E}_T$ . To remove such common balanced multijet events which gained  $\not{E}_T$ , we required that the charged tracks associated with the central jet of highest  $E_T$  be consistent with that emanating from the primary interaction vertex (80% efficient for signal) [6]. The 15 events passing this criterion formed our penultimate event sample.

The background estimate for our nominal cut set  $(\not\!\!\!E_T > 75 \text{ GeV}, H_T > 100 \text{ GeV})$  is shown in the second line of Table I. The largest noninstrumental background arises from the production of  $t\bar{t}$  pairs in which one t quark decays into jets and the other decays into  $b\ell\nu$ , where  $\ell = e, \mu$ , or  $\tau$ , and the lepton is not detected. We generated  $t\bar{t} \rightarrow b\ell\nu + \text{jets}$  events using the HERWIG Monte Carlo [7], a GEANT [8] based detailed detector simulation, and the same reconstruction program used for data. We

TABLE I. Optimized  $\not{E}_T$  and  $H_T$  thresholds for several regions of mSUGRA parameter space. The optimal thresholds were chosen for the specified  $m_0$  and  $m_{1/2}$  values that correspond to the listed gluino and squark masses. The next-to-leading-order cross sections and the total efficiency for signal events, with their combined statistical and systematic uncertainties, the total number of events expected from backgrounds, with their statistical and systematic uncertainties, the number of observed events, the probability for observing  $N_{obs}$  events or greater given the background prediction, and the 95% confidence level upper limit on the cross section for the particular ( $m_0, m_{1/2}$ ) point are given in the remaining columns. Note that the entries in this table are strongly correlated.

₿ <sub>Tthresh</sub> (GeV)	$H_{T_{\mathrm{thresh}}}$ (GeV)	$(m_0, m_{1/2})$ (GeV/ $c^2$ )	$(m_{\tilde{g}}, m_{\tilde{q}}) \ ({ m GeV}/c^2)$	$\sigma_{ m sig}$ (pb)	$\epsilon$ (%)	$N_{ m bck-pred}$	$N_{ m obs}$	P <sub>over</sub> (%)	$\sigma_{95}$ (pb)
50	100	relaxe	ed $\not\!\!\!E_T$ threshold			$43.0\pm0.8^{+8.5}_{-8.2}$	49	29.5	
75	100	(150, 80)	(243, 249)	4.4	$5.8 \pm 0.5^{+1.7}_{-1.4}$	$8.3 \pm 0.8^{+3.4}_{-3.2}$	15	9.2	4.4
75	120	(300, 50)	(172, 318)	15.7	$1.5 \pm 0.3^{+0.3}_{-0.2}$	$5.5 \pm 0.5^{+2.7}_{-2.6}$	12	6.2	14.8
75	140	(200, 80)	(246, 278)	2.4	$5.8 \pm 0.4^{+1.0}_{-1.6}$	$3.6 \pm 0.2 \pm 2.1$	11	2.0	5.1
75	150	(250, 60)	(198, 286)	7.1	$3.1 \pm 0.3^{+0.4}_{-0.9}$	$3.0 \pm 0.1 \pm 1.9$	8	6.1	8.1
75	160	(300, 70)	(228, 339)	2.0	$4.2\pm0.4^{+0.7}_{-0.8}$	$2.6 \pm 0.1^{+1.8}_{-1.7}$	6	12.9	3.3
90	100	(100, 100)	(290, 266)	1.8	$7.7 \pm 0.5^{+1.4}_{-1.5}$	$6.0 \pm 0.7^{+2.7}_{-2.5}$	8	31.8	1.7
100	100	(0, 100)	(288, 250)	2.8	$4.9 \pm 0.4^{+1.0}_{-1.1}$	$4.6\pm0.7^{+2.2}_{-2.0}$	7	25.4	2.7
100	150	(200, 110)	(322, 330)	0.3	$9.2 \pm 0.5^{+0.6}_{-1.3}$	$1.3 \pm 0.1 \pm 1.2$	3	24.4	0.9

assumed the  $t\bar{t}$  production cross section of 5.9  $\pm$  1.6 pb [9], which yielded a prediction of 3.1  $\pm$  0.2(stat)<sup>+1.4</sup><sub>-1.3</sub>(syst) background events.

Comparable backgrounds come from the production of W and Z bosons. Substantial  $\not{E}_T$  can arise in events with a W boson decaying to leptons where the charged lepton is not identified, and in events with  $Z \rightarrow \nu\nu$  or  $Z \rightarrow \tau\tau$  decays. To estimate these backgrounds, we generated Monte Carlo samples for W boson events with VECBOS [10] (quark hadronization simulated using ISAJET [11]), Z bosons with PYTHIA [12], and WW and WZ events with ISAJET. The detector response was modeled as for the  $t\bar{t}$  sample. From all vector boson production sources, we predict  $2.8 \pm 0.8^{+0.7}_{-0.5}$  events, 85% of which are from  $W \rightarrow \ell \nu$  and  $Z \rightarrow \nu \nu$  decays.

The only remaining background is events that have  $\not\!\!E_T$ because one or more jets are mismeasured. To determine this background, we used events from 56  $pb^{-1}$  of data collected with a trigger requiring at least one jet with  $E_T > 85$  GeV. The trigger was fully efficient for events containing a jet with  $E_T > 115$  GeV. Events with  $\not\!\!\!E_T < 50 \text{ GeV}$  were used to determine this instrumental background to events with larger  $\not\!\!E_T$  using two different estimations. The primary method relied on a Bayesian shape analysis [13]. We define the quantity  $D_{\pi\pi} = \sqrt{(\delta \phi_1 - \pi)^2 + (\delta \phi_2 - \pi)^2}$ , which has a distribution that is strongly peaked at large  $D_{\pi\pi}$  for events with apparent  $E_T$  due to mismeasured jets and is nearly independent of the  $\not\!\!E_T$  threshold. For  $t\bar{t}$  and signal the distribution is less peaked, as shown in Fig. 1. To determine the multijet contribution, we performed a three-component  $(t\bar{t}, multi$ jet, and signal) fit to the shape of the  $D_{\pi\pi}$  distribution in the data. The backgrounds quoted in Table I include the multijet contribution, as determined in this fit. As a check, we fit the  $\not\!\!E_T$  spectrum of our event sample between 25 and 50 GeV to an exponential in  $\not\!\!E_T$ ; extrapolation to higher  $\not\!\!E_T$ 

yielded a prediction in agreement with the fit to  $D_{\pi\pi}$ , as shown in Table II.

The final selection criteria for each  $(m_0, m_{1/2})$  point were determined by choosing  $H_T$  and  $\not\!\!E_T$  thresholds that maximized the  $S/\delta B$  ratio, where S is the expected number of SUSY events and  $\delta B$  is the combined systematic



FIG. 1. Sample  $D_{\pi\pi}$  distributions used in the Bayesian shape fitter. Note that the bins do not have uniform widths. The top plot shows data passing the analysis requirements with  $\not\!\!E_T > 75$  GeV and  $H_T > 150$  GeV (eight candidate events are accepted). The lower plot shows the  $D_{\pi\pi}$  distributions for QCD multijet (thick line),  $t\bar{t}$  (dashed line), and an mSUGRA sample with  $m_0 = 250$  GeV/ $c^2$  and  $m_{1/2} = 60$  GeV/ $c^2$  (thin line) for events passing the same requirements. The normalizations for the QCD multijet,  $t\bar{t}$ , and mSUGRA plots are 55.9, 7000, and 350 pb<sup>-1</sup>, respectively.

$E_{T_{\rm thresh}}$ (GeV)	$H_{T_{\mathrm{thresh}}}$ (GeV)	Bayesian fit to $D_{\pi\pi}$	Extrapolation
75	100	$2.5 \pm 2.6$	$2.8 \pm 0.9$
75	150	$0.8 \pm 1.6$	$1.7 \pm 0.3$
100	100	$0.7 \pm 1.6$	$0.6 \pm 0.1$

and statistical uncertainty on the background predicted from the SM. Table I shows the thresholds used. Variation of the jet energy scale [14] dominates the systematic uncertainty for Monte Carlo based background estimates.

We note that, for all of the entries in Table I, the number of observed events is greater than the number predicted from background. The results are highly correlated, since most rows are subsets of previous rows. The probability of obtaining at least the number of events observed for any of the listed cutoffs is more than 2%, and we therefore interpret our result as a constraint on the  $m_0$ and  $m_{1/2}$  parameters of mSUGRA. By simulating squark and gluino production and decay with ISAJET, followed by the same detector response and event reconstruction as in our previous simulations, we generated samples at several values of  $m_0$  and  $m_{1/2}$ , all with the mSUGRA parameters  $\tan\beta = 2$ ,  $A_0 = 0$ , and  $\mu < 0$ . Using the next-to-



FIG. 2. The exclusion contour obtained in this analysis (heavy line), the region below which is excluded at the 95% confidence level. The thin lines are contours of constant squark or gluino mass in the  $m_0 - m_{1/2}$  plane, as indicated. In the shaded region, mSUGRA does not contain electroweak symmetry breaking and is excluded *a priori*.

leading-order squark and gluino production cross sections from PROSPINO [15], and a Bayesian technique with a flat prior for the signal, we determined 95% confidence level limits on the parameters. For each  $\not{E}_T$ ,  $H_T$  threshold pair, Table I displays results from an mSUGRA signal sample. The signal efficiencies and the structure of the resulting limit contour reflect the complex variation of jet multiplicity,  $\not{E}_T$  and jet  $E_T$  spectra with changing higgsino and wino/zino content in the neutralinos and charginos for each signal sample.

Figure 2 shows the region excluded by this analysis. We exclude all mSUGRA models with  $m_{\tilde{q}} < 250 \text{ GeV}/c^2$ . For small  $m_0$  and for  $m_{\tilde{q}} = m_{\tilde{g}}$ , we exclude  $m_{\tilde{g}} < 300 \text{ GeV}/c^2$  and masses less than 260 GeV/ $c^2$ , respectively. In Fig. 3, we show the exclusion contour in the  $(m_{\tilde{g}}, m_{\tilde{q}})$  plane compared to results from other experiments. We extend significantly the limits on squarks and gluinos, especially in the region where  $m_{\tilde{g}} > m_{\tilde{q}}$ . Within mSUGRA models, for negative  $\mu$  and  $\tan\beta = 2$ , the CERN LEP limits on charginos  $m_{\tilde{\chi}_1^\pm} < 86$  to 45 GeV/ $c^2$  for small  $m_0$  and 86 GeV/ $c^2$  for large  $m_0$ . Our limit on  $m_{1/2}$ 



FIG. 3. The limit from this analysis in the  $(m_{\tilde{g}}, m_{\tilde{q}})$  mass plane ("this experiment"). The figure also shows curves of previous limits from jets and  $E_T$  from D0 [16,17] (hatched) using 7.2 pb<sup>-1</sup> of data and MSSM parameters tan $\beta = 2$  and  $\mu = -250 \text{ GeV}/c^2$ , the jets and  $\not\!\!E_T$  limit from CDF [18] based on 19 pb<sup>-1</sup> of data with  $\tan\beta = 4$  and  $\mu = -400 \text{ GeV}/c^2$ (thick dots), the dilepton CDF limit [19] (dashed-dotted line) from 19 pb<sup>-1</sup> of data with mSUGRA parameters  $\tan\beta = 4$ and  $\mu < 0$ , and limits using only direct decays from UA1/ UA2 [20] (dotted line) and DELPHI/Mark II [21] (dotted line). The MSSM analyses shown here have further assumptions explained in the references. The limits are insensitive to the small differences in mSUGRA and MSSM models used in the CDF and D0 analyses. More recent model-dependent CERN LEP limits are given in the text. All limits are at the 95% confidence level. The region below the diagonal dashed line is excluded because there the squark is the LSP.

ranges between 100 GeV/ $c^2$  for small  $m_0$  and 60 GeV/ $c^2$  for large  $m_0$ .

In summary, we have searched for events with large  $\not\!\!E_T$ and multiple jets, and observe no statistically significant excess of events beyond expectations from SM processes. This null result is interpreted in the context of minimal low-energy supergravity as an excluded region in the  $(m_0, m_{1/2})$  plane and is most pertinent in increasing the mass limits on squarks and gluinos.

We thank the Fermilab and collaborating institution staffs for contributions to this work, and we acknowledge support from the Department of Energy and National Science Foundation (USA), Commissariat à L'Energie Atomique (France), Ministry for Science and Technology and Ministry for Atomic Energy (Russia), CAPES and CNPq (Brazil), Departments of Atomic Energy and Science and Education (India), Colciencias (Colombia), CONACyT (Mexico), Ministry of Education and KOSEF (Korea), and CONICET and UBACyT (Argentina).

\*Visitor from Universidad San Francisco de Quito, Quito, Ecuador.

<sup>†</sup>Visitor from IHEP, Beijing, China.

- Yu. A. Golfand and E. P. Likhtman, JETP Lett. 13, 323 (1971); D. V. Volkov and V. P. Akulov, Phys. Lett. 46B, 109 (1973); J. Wess and B. Zumino, Nucl. Phys. B70, 39 (1974).
- [2] E. Witten, Nucl. Phys. B188, 513 (1981); R. K. Kaul and P. Majumdar, Nucl. Phys. B199, 36 (1982); J. Polchinski and L. Susskind, Phys. Rev. D 26, 3661 (1982).
- [3] G. L. Kane *et al.*, Phys. Rev. D **49**, 6173 (1994); H. Baer and X. Tata, Phys. Rev. D **47**, 2739 (1993); M. Drees and M. M. Nojiri, Nucl. Phys. **B369**, 54 (1992); L. E. Ibañez, C. Lopez, and C. Munoz, Nucl. Phys. **B256**, 218 (1985).
- [4] D0 Collaboration, S. Abachi *et al.*, Nucl. Instrum. Methods Phys. Res., Sect. A **338**, 185 (1994).
- [5] D0 Collaboration, S. Abachi *et al.*, Phys. Lett. B **357**, 500 (1995).

- [6] A. L. Lyon, Ph.D. thesis, University of Maryland, 1997 (unpublished) http://www-d0.fnal.gov/results/ publications\_talks/thesis/thesis.html
- [7] G. Marchesini *et al.*, Comput. Phys. Commun. **67**, 465 (1992).
- [8] R. Brun and F. Carminati, CERN Program Library Long Writeup W5013, 1993 (unpublished). We used GEANT V3.14.
- [9] D0 Collaboration, B. Abbott *et al.*, Phys. Rev. Lett. **83**, 1908 (1999).
- [10] F.S. Berends, H. Kuijf, H. Tausk, and W.T. Giele, Nucl. Phys. B357, 32 (1991).
- [11] F. Paige and S. Protopopescu, in Supercollider Physics, edited by D. Soper (World Scientific, Singapore, 1986), p. 41; H. Baer et al., in Proceedings of the Workshop on Physics at Current Accelerators and Supercolliders, edited by J. Hewett et al. (Argonne National Laboratory, Argonne, IL, 1993). We used ISAJET V7.13.
- [12] T. Sjöstrand, CERN-TH 6488/92 (unpublished). We used PYTHIA V5.7 with JETSET V7.4.
- [13] P.C. Bhat, H.B. Prosper, and S.S. Snyder, Phys. Lett. B 407, 73 (1997).
- [14] D0 Collaboration, B. Abbott *et al.*, Nucl. Instrum. Methods Phys. Res., Sect. A **424**, 352 (1999).
- [15] W. Beenakker *et al.*, Nucl. Phys. **B492**, 51 (1997);
   W. Beenakker *et al.*, hep-ph/9611232 (unpublished).
- [16] D0 Collaboration, S. Abachi *et al.*, Phys. Rev. Lett. **75**, 618 (1995).
- [17] D. Claes, in Proceedings of the 10th Topical Workshop on Proton-Antiproton Collider Physics, edited by R. Raja and J. Yoh, AIP Conf. Proc. No. 357 (AIP, New York, 1996).
- [18] CDF Collaboration, F. Abe *et al.*, Phys. Rev. D 56, 1357 (1997).
- [19] CDF Collaboration, F. Abe *et al.*, Phys. Rev. Lett. **76**, 2006 (1996).
- [20] UA1 Collaboration, C. Albajar *et al.*, Phys. Lett. B **198**, 261 (1987); UA2 Collaboration, J. Alitti *et al.*, Phys. Lett. B **235**, 363 (1990).
- [21] Mark II Collaboration, T. Barklow *et al.*, Phys. Rev. Lett. 64, 2984 (1990); DELPHI Collaboration, P. Abreu *et al.*, Phys. Lett. B 247, 148 (1990).
- [22] D. Haidt et al., Eur. Phys. J. C 3, 1 (1998).